CME propagation: where does aerodynamic drag "take over"?

Prasad Subramanian

Indian Institute of Science Education and Research (IISER), Pune, India

- Nishtha Sachdeva (PhD student, IISER Pune)
- Robin Colannino (Naval Research Lab, Washington, DC)
- Angelos Vourlidas (Applied Physics Lab/Johns Hopkins U)
- Volker Bothmer (U Goettingen)

• CMEs are subjected to driving and drag forces from initiation through Sun-Earth propagation

- CMEs are subjected to driving and drag forces from initiation through Sun-Earth propagation
- The broad consensus is:

- CMEs are subjected to driving and drag forces from initiation through Sun-Earth propagation
- The broad consensus is:
 - Driving some sort of Lorentz $(\mathbf{J} \times \mathbf{B})$ force,

- CMEs are subjected to driving and drag forces from initiation through Sun-Earth propagation
- The broad consensus is:
 - Driving some sort of Lorentz $(\mathbf{J} \times \mathbf{B})$ force,
 - Drag aerodynamic drag, due to "friction" of the CME with the surrounding solar wind

- CMEs are subjected to driving and drag forces from initiation through Sun-Earth propagation
- The broad consensus is:
 - Driving some sort of Lorentz $(\mathbf{J} \times \mathbf{B})$ force,
 - Drag aerodynamic drag, due to "friction" of the CME with the surrounding solar wind
- The balance between driving and drag determines the CME trajectory, and important takeaways such as Earth arrival time, speed, etc.

 $\bullet\,$ Fast CMEs (> 1000 km/s) are decelerating even as early as when they appear in a coronograph field of view -

- $\bullet\,$ Fast CMEs ($>1000\,$ km/s) are decelerating even as early as when they appear in a coronograph field of view -
- so only drag matters from there on;

- Fast CMEs (> 1000 km/s) are decelerating even as early as when they appear in a coronograph field of view -
- so only drag matters from there on;
- i.e., For the most part, fast CMEs are only dragged down by the ambient solar wind no driving

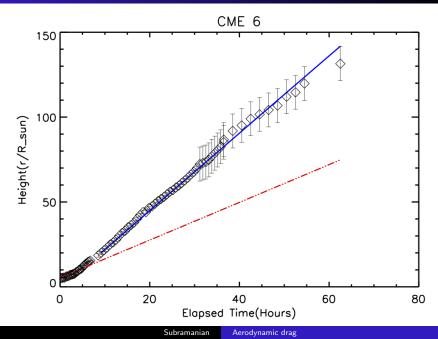
- Fast CMEs (> 1000 km/s) are decelerating even as early as when they appear in a coronograph field of view -
- so only drag matters from there on;
- i.e., For the most part, fast CMEs are only dragged down by the ambient solar wind no driving
- Nicely confirmed using physical drag presciption (not ad-hoc constant C_D!) Subramanian, Lara & Borgazzi 2012 GRL

• Conversely, very slow CMEs (a few 100 km/s) are presumed to be dragged up by the solar wind from early on (a few R_{\odot} onwards

- \bullet Conversely, very slow CMEs (a few 100 km/s) are presumed to be dragged up by the solar wind from early on (a few R_{\odot} onwards
- Not so..slow CMEs (starting speeds 100-200 km/s) modeled with a drag-only prescription (physical drag model of Subramanian et al 2012 as well as constant C_D) disagree considerably with observations..

- \bullet Conversely, very slow CMEs (a few 100 km/s) are presumed to be dragged up by the solar wind from early on (a few R_{\odot} onwards
- Not so..slow CMEs (starting speeds 100-200 km/s) modeled with a drag-only prescription (physical drag model of Subramanian et al 2012 as well as constant C_D) disagree considerably with observations..
- ...when initiated from the first timestamp (as is usually done)

Drag only model fails when initiated from start



• Drag-only models agree with observations when they are initiated farther out; i.e., not from the first timestamp

- Drag-only models agree with observations when they are initiated farther out; i.e., not from the first timestamp
- i.e., slow(er) CMEs are drag dominated only beyond 15–50
 *R*_☉ (Sachdeva et al 2015 ApJ)

- Drag-only models agree with observations when they are initiated farther out; i.e., not from the first timestamp
- i.e., slow(er) CMEs are drag dominated only beyond 15–50
 *R*_☉ (Sachdeva et al 2015 ApJ)
- So perhaps Lorentz forces are important until pprox 15–50 R_{\odot} ?

- Drag-only models agree with observations when they are initiated farther out; i.e., not from the first timestamp
- i.e., slow(er) CMEs are drag dominated only beyond 15–50 R_{\odot} (Sachdeva et al 2015 ApJ)
- So perhaps Lorentz forces are important until \approx 15–50 $R_{\odot}?$

No.	CME date	<i>v</i> _0(km/s)	$\widetilde{h}_0(R_\odot)$
1	Mar 19-23,2010	162	21.9
2	Apr 03-05,2010	916	5.5
3	Apr 08-11,2010	468	19.7
4	Jun 16-20,2010	193	15.2
5	Sept 11-14,2010	444	27.7
6	Oct 26-31,2010	215	20.1
7	Feb 15-18,2011	832	39.7
8	Mar 25-29,2011	47	46.5

Sachdeva et al 2015 ApJ

$$\rho_m \frac{d^2 R}{dt^2} = \frac{I^2}{4\pi^2 b^2 R^2} (L + \mu_0 R/2) - \frac{IB_{\text{ex}}(R)}{\pi b^2}$$

 ..following Kliem & Torok (2006) - includes both Lorentz self-forces and Lorentz external forces

$$\rho_m \frac{d^2 R}{dt^2} = \frac{I^2}{4\pi^2 b^2 R^2} (L + \mu_0 R/2) - \frac{IB_{ex}(R)}{\pi b^2}$$

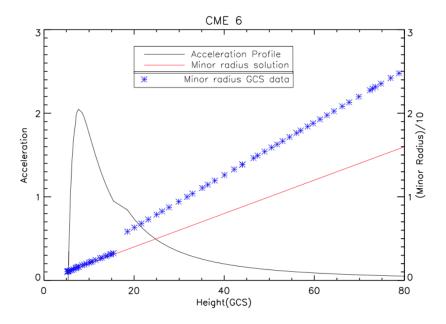
- ..following Kliem & Torok (2006) includes both Lorentz self-forces and Lorentz external forces
- If external (constraining) field falls off fast enough, there will be a "whiplash" the toroidal instability, which flings/launches the CME

$$\rho_m \frac{d^2 R}{dt^2} = \frac{I^2}{4\pi^2 b^2 R^2} (L + \mu_0 R/2) - \frac{IB_{ex}(R)}{\pi b^2}$$

- ..following Kliem & Torok (2006) includes both Lorentz self-forces and Lorentz external forces
- If external (constraining) field falls off fast enough, there will be a "whiplash" - the toroidal instability, which flings/launches the CME
- The instability (well known in lab plasmas) works if $B_{ex}(R) \propto R^{-n}$, n > 3/2

$$\rho_m \frac{d^2 R}{dt^2} = \frac{I^2}{4\pi^2 b^2 R^2} (L + \mu_0 R/2) - \frac{IB_{ex}(R)}{\pi b^2}$$

- ..following Kliem & Torok (2006) includes both Lorentz self-forces and Lorentz external forces
- If external (constraining) field falls off fast enough, there will be a "whiplash" - the toroidal instability, which flings/launches the CME
- The instability (well known in lab plasmas) works if $B_{ex}(R) \propto R^{-n}$, n > 3/2
- Naturally predicts a peak in the Lorentz force; i.e., force dies down at large *R*



 ...using a model that includes *only* Lorentz forces (no drag), we find

Looking forward...

- ...using a model that includes *only* Lorentz forces (no drag), we find
- Lorentz forces cease to be dominant at $\approx 20R_{\odot}$;

Looking forward...

- ...using a model that includes *only* Lorentz forces (no drag), we find
- Lorentz forces cease to be dominant at $\approx 20R_{\odot}$;
- ...just where our drag-only model predicted that drag forces would start to dominate

- ...using a model that includes *only* Lorentz forces (no drag), we find
- Lorentz forces cease to be dominant at $\approx 20R_{\odot}$;
- ...just where our drag-only model predicted that drag forces would start to dominate
- Even the minor radius expansion (for which data is much better) agrees very well with this finding

- ...using a model that includes *only* Lorentz forces (no drag), we find
- Lorentz forces cease to be dominant at $\approx 20R_{\odot}$;
- ...just where our drag-only model predicted that drag forces would start to dominate
- Even the minor radius expansion (for which data is much better) agrees very well with this finding
- ...so we are progressing towards a physically motivated understanding (not simply parametrization/fitting) of drag, as well as drive forces acting on CMEs

- ...using a model that includes *only* Lorentz forces (no drag), we find
- Lorentz forces cease to be dominant at $\approx 20R_{\odot}$;
- ...just where our drag-only model predicted that drag forces would start to dominate
- Even the minor radius expansion (for which data is much better) agrees very well with this finding
- ...so we are progressing towards a physically motivated understanding (not simply parametrization/fitting) of drag, as well as drive forces acting on CMEs
- Crucial for reliable time-of-arrival estimates
- Thank you for your attention!